



## Truth, its theory and its practice in science education

### *ABSTRACT:*

This presentation argues that science education must embrace truth.

It draws upon the origins of modern science, particularly the work of Isaac Newton on optics, to show how truth is essential to science.

Newton wrestles with truth both in his early experiments and later as he attempts to present his work to a sceptical audience. The formulation of truth in this presentation is that of Martin Heidegger.

If we are to make science understandable to students, and inspire them, we must present them with an adequate account of science itself, and this can only be achieved if we embrace a concept of truth.

# Today



Intellectual tradition

Truth

Heidegger's theory

Discovery as truth

Truth in modern science

Heidegger's characteristics

Newton

Implications

Philosophy of science

Science education

# Hermeneutic philosophy of science

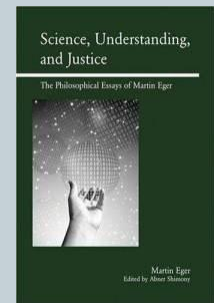
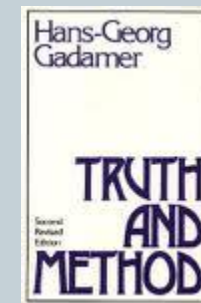
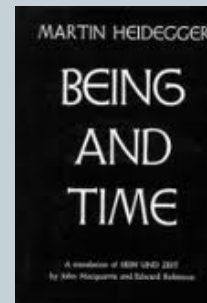


Without truth

Heidegger

Gadamer

Eger



With truth

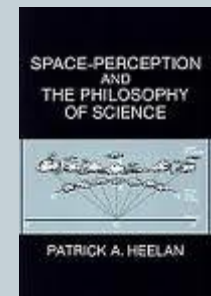
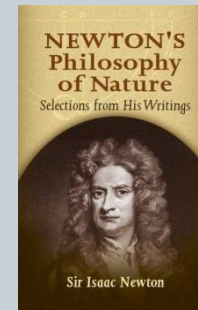
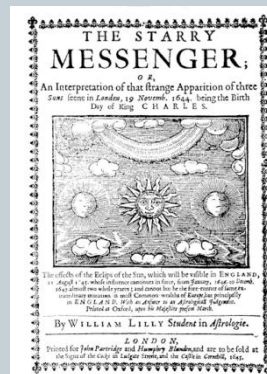
Galileo 1610

Newton 1666

Kant 1782

Heidegger 1927

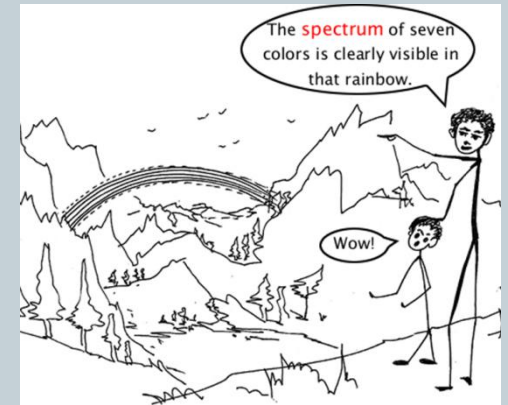
Heelan 1983



# Heidegger's concept of truth



Correspondence



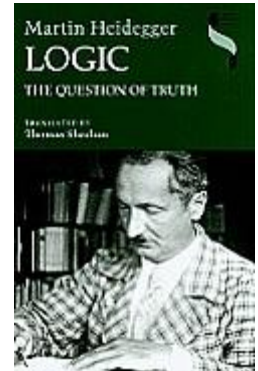
Disclosure







# Three characteristics of modern science



Framework / ground plan  
Pre-logic  
Perception  
Measurable objects  
The Real

Force nature to reveal more of itself  
Follow procedures  
Equipment  
Event of truth  
Individual  
Perception

Consequences of disclosures  
Literature  
Institutions  
Workforce



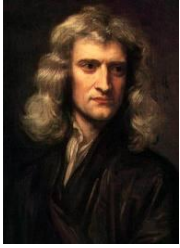






Обращения Девиак  
1610

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2. янв.	○ ** *
3. мар'	○ * *
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10. марк.	* * * ○ *
11.	* * ○ *
12. H. 4 уел.	* ○ *
17. мар'	* ** ○ *
14. янв.	* * * ○ *



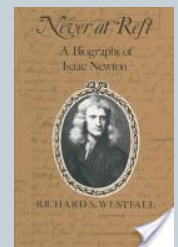
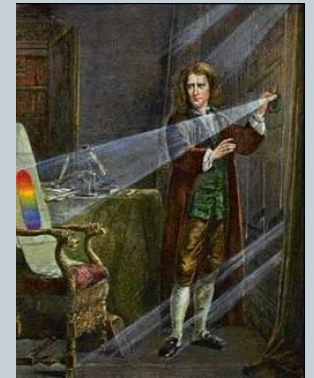
# Newton's optics

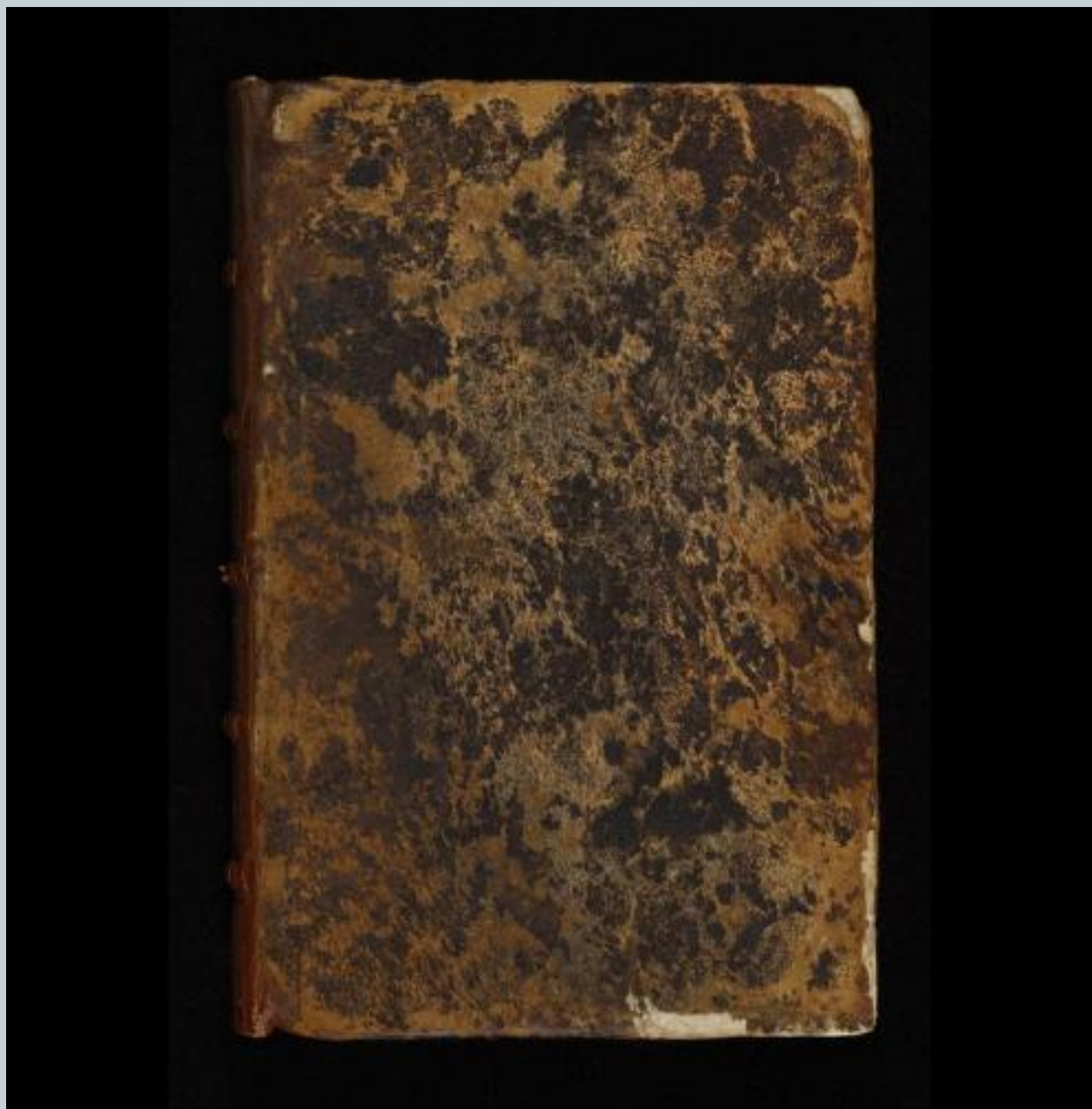


1664  
22 years-old  
Trinity College

*Questiones quædam Philosophiæ*  
Certain philosophical questions

*Amicus Plato amicus Aristoteles magis amica veritas*  
Plato and Aristotle are my friends, but truth is a better friend





# Questions quædam Philosophæ Off. y<sup>i</sup> first matter

1

88

Whether it be mathematicall points: or Mathe-  
maticall points & parts: or a simple entity  
before division indistinct: or individuall *ie.* Atoms

Not of Mathematicall points since *at* wants di-  
visions cannot substitute a body in thire  
conjunction because they will sink into y<sup>i</sup>  
same point. An infinite number of mathe-  
maticall points sink into one being added together  
as y<sup>t</sup> being still a mathematicall point is indivi-  
sible but a body is divisible. *et* first a Mathe-  
maticall point is Nothing since it is but an  
imaginary entity. *et*

Not of parts & Math: points, for such a point is  
either something or nothing. if something tis a pt  
& so added betwixt 2 pts will make a line  
of 3 pts. if nothing, then added betwixt two  
parts there is still nothing betwixt y<sup>e</sup> 2 pts &  
consequently y<sup>e</sup> line consists of nothing still but  
2 pts. *et*

Not of simple entity before division indistinct for tis  
not be an union of y<sup>e</sup> parts into with a body is  
difficult since these parts may againe be united as  
become one body as they were before <sup>at</sup> *ie.* nature  
of union (being but a medall one) is to depend on  
its pts (well are absolute entities) therefore it  
cannot be y<sup>e</sup> term of creation, or first matter  
for tis a contradiction to say y<sup>e</sup> first matter de-  
pends on some other subject ~~for~~ since  
y<sup>e</sup> implies some former matter or well it must  
depend.

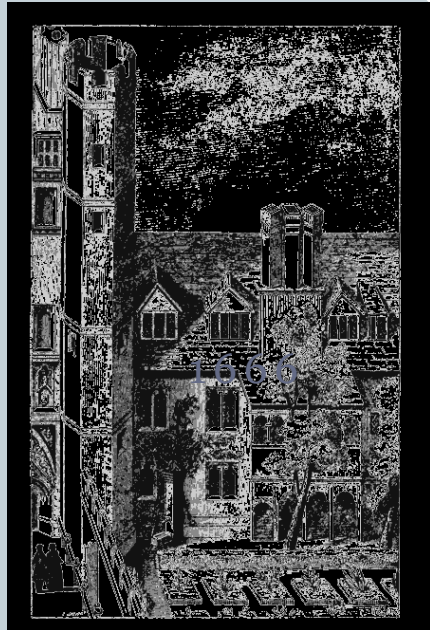
# Newton's optics

## Truth as correspondence

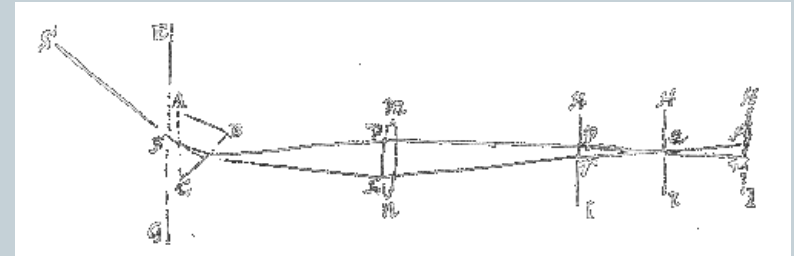
1666



white	& cōs	blew	y <sup>h</sup>	yeilo Red.
black		blew		Greene blew.
blew		black		Greene, or Red
black		red		blew.
red		black		redder.
If abdc be red	& cōs be white	y <sup>h</sup> eodc is	blew.	
white		red		redder.
white		whiter		blew.
whiter		white		redd.
black		blackcr		Greene or darke red
blackcr		black		blew.



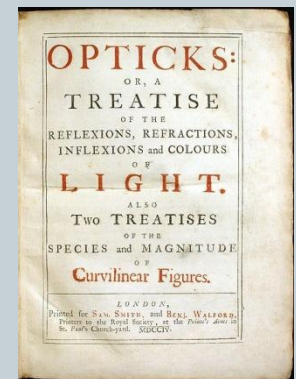
1667  
Fellow at Trinity  
College



1672 Royal Society



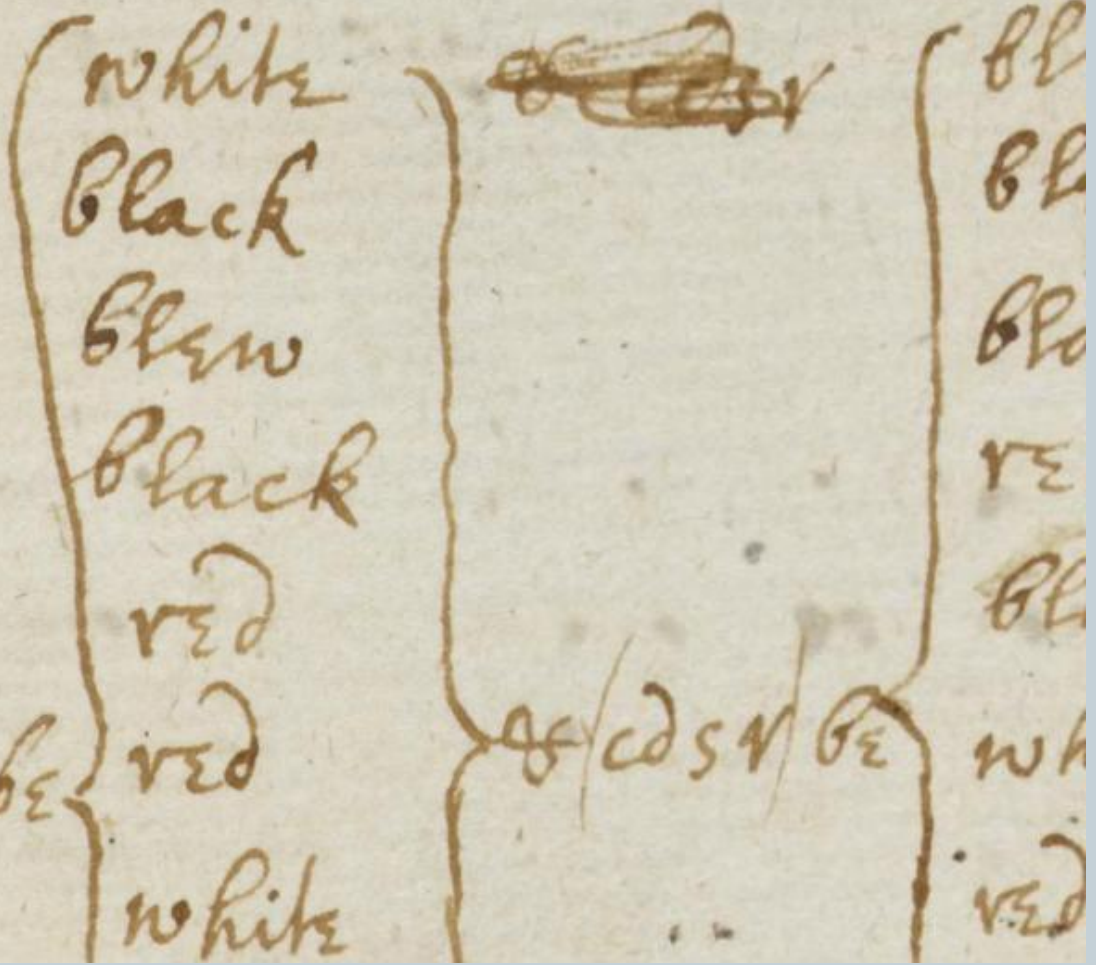
1692



1704

akz  
finding  
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If abdc be blew & cdsv whi

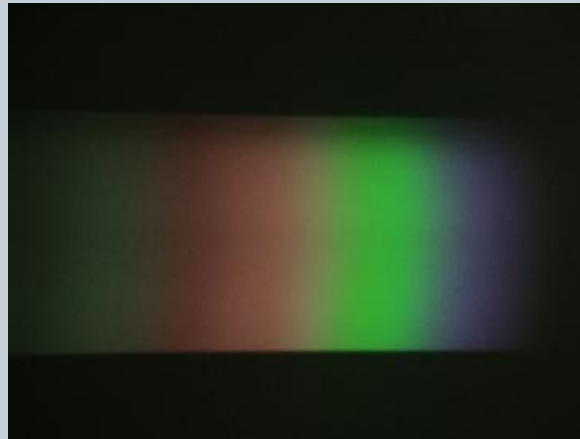


If (abdc) be

bl  
bl  
bl  
re  
bl  
wh  
red

# Newton's optics

## Truth as disclosure



# Implications - philosophy of science



Positivists & constructivists

Relativism

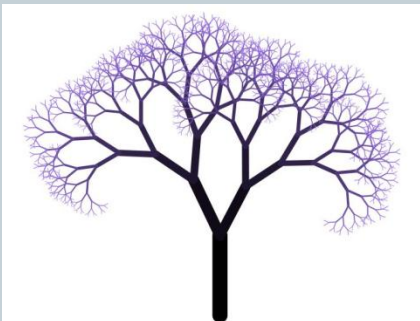
Technology

Perpetuation of science

Biology

DNA / Cold Spring Harbor

Mandelbrot's fractals



$$\begin{aligned} \mathcal{L}_{\text{StandardModel}} = & -\frac{1}{2}\partial_\mu g_\nu^\alpha \partial_\mu g_\nu^\alpha - g_\nu f^{abc} \partial_\mu g_\nu^a g_\mu^b g_\mu^c - \frac{1}{4}g_\nu^2 f^{abc} f^{ade} g_\nu^b g_\nu^c g_\nu^d g_\nu^e + \\ & \frac{1}{2}ig^2(\bar{\psi}^\mu \gamma^\mu \psi^\mu)g_\nu^\mu + G^{\mu\nu}G^{\mu\nu} + g_\nu f^{abc} \partial_\mu G^{\mu\nu} G^{\mu\nu} - \partial_\mu W_\nu^+ \partial_\mu W_\nu^- - \\ & M^2 W_\nu^+ W_\nu^- - \frac{1}{2}\partial_\mu Z_\nu^0 \partial_\mu Z_\nu^0 - \frac{1}{2}M^2 Z_\nu^0 Z_\nu^0 - \frac{1}{2}\partial_\mu A_\nu \partial_\mu A_\nu - \frac{1}{2}\partial_\mu H \partial_\mu H - \\ & \frac{1}{2}m_h^2 H^2 - \partial_\mu \phi^+ \partial_\mu \phi^- - M^2 \phi^+ \phi^- - \frac{1}{2}\partial_\mu \phi^0 \partial_\mu \phi^0 - \frac{1}{2}M^2 \phi^0 \phi^0 - \beta_h \frac{2\lambda H^2}{g^2} + \\ & \frac{2\lambda H}{g} + \frac{1}{2}(H^2 + \phi^0 \phi^0 + 2\phi^+ \phi^-) + \frac{2\lambda H^4}{g} \alpha_h - ig_{cw}[\partial_\mu Z_\nu^0(W_\nu^+ W_\nu^- - \\ & W_\nu^+ W_\nu^-) - Z_\nu^0(W_\nu^+ \partial_\mu W_\nu^- - W_\nu^- \partial_\mu W_\nu^+) + Z_\nu^0(W_\nu^- \partial_\mu W_\nu^+ - \\ & W_\nu^+ \partial_\mu W_\nu^-)] - ig_{sw}[\partial_\mu(A_\nu(W_\nu^+ W_\nu^- - W_\nu^- W_\nu^+) - A_\nu(W_\nu^+ \partial_\mu W_\nu^- - \\ & W_\nu^- \partial_\mu W_\nu^+) + A_\nu(W_\nu^+ \partial_\mu W_\nu^- - W_\nu^- \partial_\mu W_\nu^+)] - \frac{1}{2}g^2 W_\nu^+ W_\nu^- W_\nu^+ W_\nu^- + \\ & \frac{1}{2}g^2 W_\nu^+ W_\nu^- W_\nu^+ W_\nu^- + g^2 c_w^2(Z_\nu^0 W_\nu^+ Z_\nu^0 W_\nu^- - Z_\nu^0 W_\nu^+ W_\nu^-) + \\ & g^2 s_w^2(A_\nu W_\nu^+ A_\nu W_\nu^- - A_\nu A_\nu W_\nu^+ W_\nu^-) + g^2 s_w c_w[A_\nu Z_\nu^0(W_\nu^+ W_\nu^- - \\ & W_\nu^- W_\nu^+) - 2A_\nu Z_\nu^0(W_\nu^+ W_\nu^-)] - g\alpha[H^3 + H\phi^0 \phi^0 + 2H\phi^+ \phi^-] - \\ & \frac{1}{2}g^2 \alpha_h [H^4 + (\phi^0)^4 + 4(\phi^+ \phi^-)^2 + 4(\phi^0)^2 \phi^+ \phi^- + 4H^2 \phi^+ \phi^- + 2(\phi^0)^2 H^2] - \\ & gMW_\nu^+ W_\nu^- H - \frac{1}{2}ig\frac{M^2}{g^2}Z_\nu^0 Z_\nu^0 H - \frac{1}{2}ig[W_\nu^+(\phi^0 \partial_\mu \phi^- - \phi^- \partial_\mu \phi^0) - \\ & W_\nu^-(\phi^0 \partial_\mu \phi^+ - \phi^+ \partial_\mu \phi^0)] + \frac{1}{2}ig[W_\nu^+(H\partial_\mu \phi^- - \phi^- \partial_\mu H) - W_\nu^-(H\partial_\mu \phi^+ - \\ & \phi^+ \partial_\mu H)] + \frac{1}{2}ig\frac{1}{g^2}(Z_\nu^0(H\partial_\mu \phi^0 - \phi^0 \partial_\mu H) - ig_{cw}^2 M Z_\nu^0(W_\nu^+ \phi^- - W_\nu^- \phi^+) + \\ & ig_{sw} M A_\nu(W_\nu^+ \phi^- - W_\nu^- \phi^+) - ig\frac{1-2c_w^2}{2c_w}Z_\nu^0(\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) + \\ & ig_{sw} A_\nu(\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) - \frac{1}{4}g^2 W_\nu^+ W_\nu^- [H^2 + (\phi^0)^2 + 2\phi^+ \phi^-] - \\ & \frac{1}{2}g^2 \frac{1}{g^2} Z_\nu^0 [H^2 + (\phi^0)^2 + 2(2s_w^2 - 1)^2 \phi^+ \phi^-] - \frac{1}{2}g^2 \frac{2c_w}{g^2} Z_\nu^0 \phi^0 (W_\nu^+ \phi^- + \\ & W_\nu^- \phi^+) - \frac{1}{2}ig^2 \frac{2c_w}{g^2} Z_\nu^0 H (W_\nu^+ \phi^- - W_\nu^- \phi^+) + \frac{1}{2}g^2 s_w A_\nu \phi^0 (W_\nu^+ \phi^- + \\ & W_\nu^- \phi^+) + \frac{1}{2}ig^2 s_w A_\nu H (W_\nu^+ \phi^- - W_\nu^- \phi^+) - g^2 \frac{2c_w}{2c_w} (2c_w^2 - 1) Z_\nu^0 A_\nu \phi^+ \phi^- - \\ & g^2 s_w^2 A_\nu A_\mu \phi^+ \phi^- - e^3(\gamma\partial + m_\Delta^2)e^\lambda - \bar{\nu}^\lambda \gamma \partial \nu^\lambda - \bar{u}_2^2(\gamma\partial + m_\Delta^2)u_2^2 - \\ & \bar{d}_2^2(\gamma\partial + m_\Delta^2)d_2^2 + ig_{sw} A_\mu [-(e^2 \gamma^\mu e^\lambda) + \frac{2}{3}(\bar{u}_2^2 \gamma^\mu u_2^2)] - \frac{1}{2}(\bar{d}_2^2 \gamma^\mu d_2^2) + \\ & \frac{ig}{4c_w} Z_\nu^0 [(\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5))\nu^\lambda + (e^2 \gamma^\mu (4s_w^2 - 1 - \gamma^5))e^\lambda] + (\bar{u}_2^2 \gamma^\mu (\frac{2}{3}s_w^2 - \\ & 1 - \gamma^5)u_2^2) + (\bar{d}_2^2 \gamma^\mu (1 - \frac{8}{3}s_w^2 - \gamma^5)d_2^2) + \frac{ig}{2\sqrt{2}} W_\nu^+ [(\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^2))e^\lambda + \\ & (\bar{u}_2^2 \gamma^\mu (1 + \gamma^2)C_{\lambda\lambda} d_2^2)] + \frac{ig}{2\sqrt{2}} W_\nu^- [(\bar{e}^\lambda \gamma^\mu (1 + \gamma^2))\nu^\lambda + (\bar{d}_2^2 C_{\lambda\lambda}^1 \gamma^\mu (1 + \\ & \gamma^5)u_2^2) + \frac{ig}{2\sqrt{2}} \frac{m_\Delta^2}{M} [-\phi^+(1 - \gamma^5)e^\lambda + \phi^-(e^\lambda (1 + \gamma^2))\nu^\lambda] - \\ & \frac{g}{2} \frac{m_\Delta^2}{M} [H(e^\lambda e^\lambda) + i\phi^0(e^\lambda \gamma^5 e^\lambda)] + \frac{ig}{2\sqrt{2}} \phi^+ [-m_\Delta^2 (\bar{u}_2^2 C_{\lambda\lambda} (1 - \gamma^5) d_2^2) + \\ & m_\Delta^2 (\bar{d}_2^2 C_{\lambda\lambda} (1 + \gamma^5) d_2^2)] + \frac{ig}{2\sqrt{2}} \phi^- [m_\Delta^2 (\bar{d}_2^2 C_{\lambda\lambda}^1 (1 + \gamma^5) u_2^2) - m_\Delta^2 (\bar{d}_2^2 C_{\lambda\lambda}^1 (1 - \\ & \gamma^5) u_2^2)] - \frac{g}{2} \frac{m_\Delta^2}{M} H (\bar{u}_2^2 u_2^2) - \frac{g}{2} \frac{m_\Delta^2}{M} H (\bar{d}_2^2 d_2^2) + \frac{ig}{2\sqrt{2}} \phi^0 (\bar{u}_2^2 \gamma^5 u_2^2) - \\ & \frac{ig}{2\sqrt{2}} \phi^0 (\bar{d}_2^2 \gamma^5 d_2^2) + \bar{X}^+ (\partial^2 - M^2) X^+ + \bar{X}^- (\partial^2 - M^2) X^- + \bar{X}^0 (\partial^2 - \\ & \frac{M^2}{2}) X^0 + \bar{Y} \partial^2 Y + ig_{cw} W_\nu^+ (\partial_\mu \bar{X}^0 X^- - \partial_\mu \bar{X}^+ X^0) + ig_{sw} W_\nu^+ (\partial_\mu \bar{Y} X^- - \\ & \partial_\mu \bar{X}^+ Y) + ig_{cw} W_\nu^- (\partial_\mu \bar{X}^- X^0 - \partial_\mu \bar{X}^0 X^+) + ig_{sw} W_\nu^- (\partial_\mu \bar{X}^- Y - \\ & \partial_\mu \bar{Y} X^+) + ig_{cw} Z_\nu^0 (\partial_\mu \bar{X}^+ X^- - \partial_\mu \bar{X}^- X^+) + ig_{sw} A_\nu (\partial_\mu \bar{X}^+ X^- - \\ & \partial_\mu \bar{X}^- X^+) - \frac{1}{2}gM[\bar{X}^+ X^+ H + \bar{X}^- X^- H + \frac{1}{2}\bar{X}^0 X^0 H] + \\ & \frac{1-2c_w^2}{2c_w} igM[\bar{X}^+ X^0 \phi^- - \bar{X}^- X^0 \phi^-] + \frac{1}{2}igM[\bar{X}^0 X^- \phi^- - \bar{X}^0 X^+ \phi^-] + \\ & igMs_w[\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^+] + \frac{1}{2}igM[\bar{X}^+ X^+ \phi^+ - \bar{X}^- X^- \phi^0]. \end{aligned}$$

# Implications - science education



Truth

Demonstrations Experiments Simulations

Hypothesis

Mathematical models

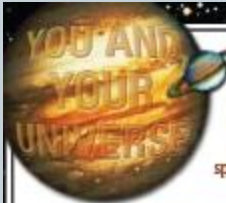
Scientific method

Curriculum structure

Science literacy

Nature of science



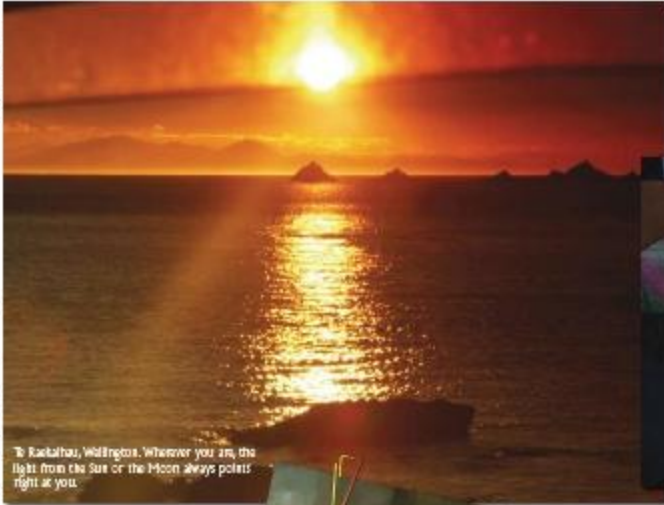


# Astronomy and you

You are you. No one else can be you. What you think and do is yours alone. You are secret within yourself. You occupy a particular space. You are not somewhere else, just in this one place.

## You are significant

You live in a particular time. You did not appear earlier on when the world was a different place. Nor will you appear later – right now is your time! You see things from where you are. Your view of everything is unique. It is not possible for anyone else to see what you see.



Te Kaitiaki, Wellington. Wherever you are, the light from the Sun or the Moon always points right at you.

## You are insignificant

Have you ever had that feeling that you are very tiny? You are in an enormous Universe. This is your brief moment in space and time. There were millions of people before you. There will be millions of people after you. No matter what you do, you cannot be significant in the long term.

Our world has not always existed. Nor will it always exist. One day everything we know will be gone. Astronomy helps us all to see ourselves in perspective.

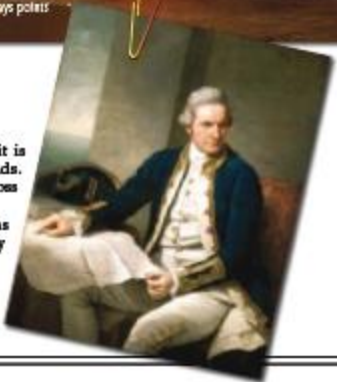
On the one hand we are important to ourselves, but on the other hand we are not important at all.

## History and culture

Astronomy is important in the history of all cultures. However, it is particularly important for people who live on small, isolated islands. New Zealanders all originated as migrants, and they had to travel across a vast ocean to get to our islands.

In their journeys to Aotearoa, both the Maori and the Europeans depended on their knowledge of astronomy. Because of this, astronomy links you to your own history – wherever you came from.

Captain James Cook (1728–1779). Captain Cook's crew relied on astronomical observations for navigation on their voyages of discovery to the Pacific, as did the Pacific peoples generations earlier. (Painting by Nathaniel Dance, 1725)



## Our future

If we are to enjoy a good standard of living, our country must earn money overseas. The only way to do this is by being a creative and intelligent nation. When you study – be it science or anything else – you make a contribution towards New Zealand's future. Science, the arts, and the humanities all offer ideas and insights that are important to our country.



The portable dome used by the Carter Observatory to teach children about space.



Using a globe and a light, a student from Palmerston North Intermediate Normal School demonstrates the reason for Earth's seasons to others in her group.

## Exciting revelations in unsettled times

More is being revealed about our Universe all the time. Things out there are not as we thought they might be. New discoveries and new ideas always produce more questions. Once you learn the basics of astronomy, you will understand the importance of each discovery.

You happen to live in an age where there is a dramatic expansion in human knowledge. Scientific understanding is undergoing major change. Astronomy is a part of this, but all the sciences are included.

Beyond science there are also great conflicts of ideas, for example the global conflicts between religious and political ideas. In other words, you live in a time that is unsettled and uncertain, rather than stable and reliable – a time when there is debate about what is true.





## New Zealand astronomers and observatories

New Zealanders have a long history of involvement in science. The first learned society, the Literary and Scientific Institution of Nelson, was founded on the emigrant ship *Whitby* whilst it was at sea in April 1841.

In 1851, English migrant Charles Roeking Carter gave the first recorded address to the New Zealand Society, which went on to become The Royal Society of New Zealand (the professional association of New Zealand scientists).

Present-day astronomy sites in New Zealand include:

- universities, particularly the University of Auckland and the University of Canterbury
- research observatories, particularly the Mt John University Observatory
- local observatories, such as the Stardome in Auckland, the R.F. Joyce Observatory West Melton, and the Beverly-Bagg Observatory in Dunedin.

The Astronomy Australia website ([www.astronomyaustralia.net](http://www.astronomyaustralia.net)) provides links to the New Zealand organisations.

### John Hearnshaw – professional astronomer

John Hearnshaw is professor of astronomy at the University of Canterbury. He teaches students and carries out research in astronomy. He has written three books on astrophysics and is New Zealand's senior astronomer.

John Hearnshaw is also active in the administration of astronomy. As a member of the International Astronomical Union (IAU), he chairs a group that advances astronomy teaching and research in developing countries.

The IAU was formed in 1919. Today it involves over 9000 astronomers from 87 countries. It is important for New Zealand that our scientists are involved in international scientific organisations, because those scientists keep everyone in New Zealand informed about new ideas.



John Hearnshaw, New Zealand senior astronomer

### Beatrice Tinsley – woman of ideas

- Born in Chester, England, in 1941
- Came to New Zealand when she was five years old
- Father was mayor of New Plymouth
- Became an eminent cosmologist (someone who studies the origin, development, and nature of the physical Universe)
- Her thesis was entitled *The Evolution of Galaxies and its Significance for Cosmology*

- Was a theoretical astronomer (worked without a telescope)
- Became Professor of Astronomy at Yale University
- Died of skin cancer aged 40 years, in 1981.

Tinsley proposed and developed the idea of galactic evolution – how galaxies are formed, and how they change over long periods of time. Her work concerned groups of stars and how they age together. She developed the idea of the ‘protogalaxy’ – a cloud of dust that forms a galaxy. She related changes in the age of galaxies to how they looked.

Later, Tinsley carried out important work on whether the Universe is open or closed – whether or not space goes on forever.

### William Pickering – peerless manager

- Born in Wellington, in 1910
- Mother died when he was six; lived at Havelock and later boarded at Wellington College
- Inspired by his mathematics teacher ‘Pop’ Gifford (see page 26)
- Studied engineering at the University of Canterbury and the California Institute of Technology (Caltech)
- Returned to New Zealand in 1932 but, unable to find work, returned to Caltech and obtained a doctorate in physics
- Taught electrical engineering
- Became a US citizen in 1941
- Worked for NASA's Jet Propulsion Laboratory (JPL) during WWII
- Worked with van Allen (see page 71) and Werner von Braun, the German rocket scientist who massacrined the V2 rockets that dominated London (see page 48)
- Appointed director of JPL in 1954
- Received many awards and honours, including the Order of New Zealand, an Honorary Fellowship of the Royal Society of New Zealand, and an honorary knighthood (1976)
- Was twice on the cover of America's *Time* magazine
- Died in 2004.



William Pickering with US President Lyndon B. Johnson

Sir William Pickering was one of the world's leading space scientists and science managers. In November 1957 the first Soviet *Sputnik* was launched (see page 49). Just 83 days later the Americans sent *Explorer 1* into space, under the direction of Pickering. This quiet, mild-mannered New Zealander was the man responsible for America's first satellite.

### Amateur astronomers

Amateur (unpaid) astronomers do work that the professionals cannot do – they constantly monitor the sky. Because the sky is so vast, and looks different around the world, amateurs are important observers.

- There are amateur astronomers hard at work all over New Zealand. They:
- build telescopes
  - observe the sky
  - collect data
  - use the internet to exchange information
  - support observatories and astronomy clubs.

Some amateurs look for small changes that occur in the brightness of stars. Stars that change their brightness are called ‘variable stars’. It is very difficult to notice changes in the brightness of stars – the changes are very small, and you need good judgement.

Many would say that Newton was the most important scientist of all time. He was what people today would call a workaholic. During his most productive years, his long hair was frequently dirty, and he avoided people for long periods of time.

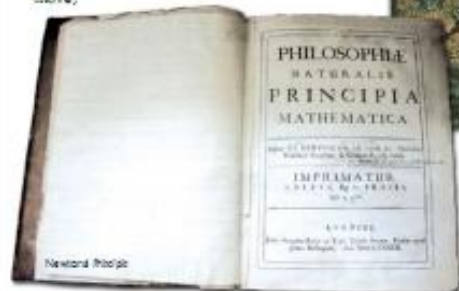
Newton was the founder of modern physics. He based his books on what he saw, and on the experiments of others. In particular he studied Copernicus and Galileo.

Some of Newton's achievements were:

- the invention of calculus (Leibniz (1646–1716) also invented calculus, probably without knowing of Newton's work, and around 225 BC the Greek Archimedes had ideas about calculus)
- a book on optics (light, colour, mirrors, and lenses)
- the theory of gravitation (see below)
- his book entitled *The Mathematical Principles of Natural Philosophy*, published in 1687. (The old name for physics was 'natural philosophy'.) This book used mathematics to describe how things move;



William Stiles' picture Newton painted in about 1685, shows the 'thinking man' – the God-like mathematician. Stiles chose to Newton's 'disordered thinking'.



- Newton's method, which is still the basic method of science. It states that in science you:
  - observe
  - measure
  - generalise (produce laws using mathematics)
  - experiment (predict results and try things out)
- set your arguments in a logical order, starting with your assumptions and definitions
- keep an open mind.

Newton believed that the laws he discovered were God's work. As he wrote (in Latin), 'God created everything by number, weight and measure'.

The important point is that Newton used mathematics to explain the things he saw; he measured things, then used mathematics to try to explain how different things were related.

Newton's *Principia*, published in 1687, is perhaps the most important book ever written. This photograph above shows Newton's copy of the first edition, and if you look closely you can see his corrections for the second edition.

Newton is famous for the idea of gravity. People say his theory was inspired while watching an apple fall off a tree around 1666. However, it is more correct to say that he worked on the theory for more than 10 years. Gravity was not a sudden 'bright idea'.

Gravity is the idea that everything (matter) in the Universe attracts everything else. Newton understood that when an apple falls off a tree, the Earth itself and the apple both pulled towards each other. Because the apple was lighter, it moved further. However, the Earth does move up a little to meet the apple.

## Immanuel Kant

Kant was born in Prussia (now Kaliningrad, Russia), in 1724. He was brought up in a religious household. He studied science, and became a university lecturer. He died in 1804.

Kant began to ask basic questions about science itself. He asked:

- Why is Newton's physics so successful?
- What is the Universe?
- What is reality?
- What is the nature of the things that exist?

Kant rejected the use of 'God' as an explanation for these things.

Kant made a famous statement which is inscribed on his tombstone in both German and Russian. Translated into English, it reads:

"Two things fill the mind with ever new and increasing admiration and awe – the starry sky above me and the moral law within me."



Electromagnetic wavelengths – those we can see, and those we can only detect using instruments (see page 44).

Kant's work in astronomy included:

- the nebula theory, about how our Sun and its planets came to form. Kant said that the solar system formed from a disk of gas and dust (called a nebula). Laplace (1749–1827) later developed this idea which is broadly correct.
- the idea that 'spiral nebulae' were galaxies beyond our own firm.
- the prediction of the existence of the planet Uranus. (The actual discovery of Uranus was made by Herschel in 1781)
- predictions about comets
- the claim that there must be intelligent life elsewhere in the Universe.
- The rules of mathematics are not certain. It is possible to have several forms (kinds) of mathematics. In Einstein's physics, for example, there is the idea that a straight line in space is not straight – that space is curved.
- We cannot ever imagine the Universe, because we cannot understand the idea of infinity.
- Human minds are only capable of certain kinds of understanding. There may be other forms of understanding that are always beyond us.
- Understanding is always based on our ability to think or reason, not on a reality that is 'out there'.

Although Kant's work in science was useful, his most important work is in philosophy. This includes work on the philosophy of science.

Some of Kant's key ideas were:

- Space and time may not really exist. They are perhaps generated within us by the action of our minds.

Perhaps Kant's most important idea was that human beings are limited by their own senses (sight, hearing, touch, smell, and taste), and by the way their minds are made and work.

Look at the diagram above. The upper part shows the wavelengths of light humans can see. The lower part shows the wavelengths that we know about only using instruments.



**Truth, its theory and its practice in science education**

*THANK YOU*

# Heidegger's phenomenology - enquire into human-ness



Ontological structure

Ontological understanding

Disposition

Nomination

Abidance's of truth

Disclosure

Correspondence

Care structure

Casting ahead

Hermeneutic circle

Equiprimordial

Potentiality

Ontological kinetics

Ways of being in the world

For-the-sake-of-which cascades